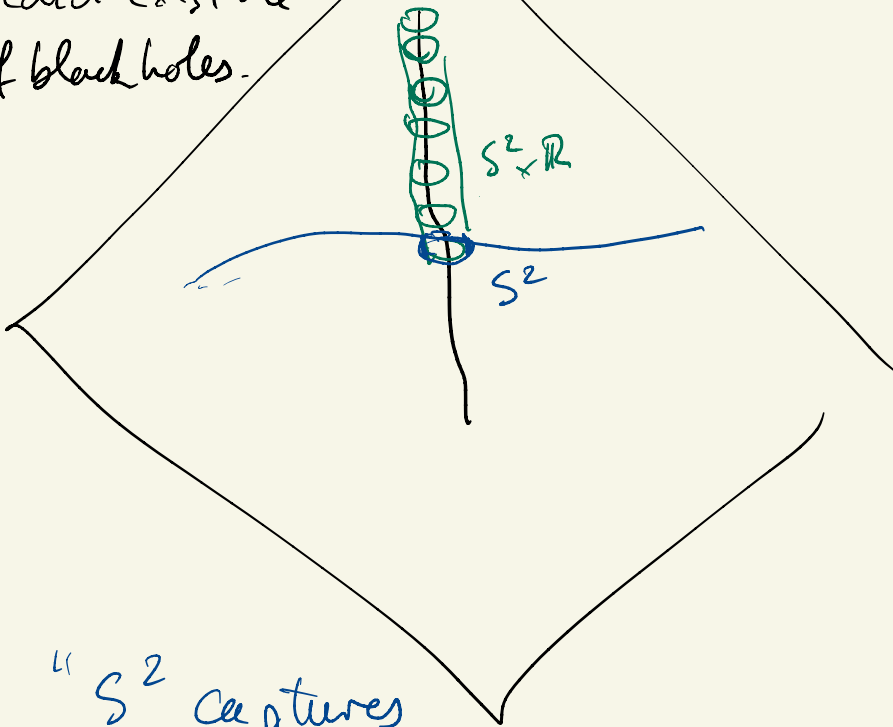


(3.) 2.9 Penrose singularity theorem

Spacetime (3+1)-dim

Predict existence of black holes.



" S^2 captures the black hole"

→ codim 2 spacelike hypersurfaces will be important

Def 2.166

A connected time-oriented
Lorentzian manifold M is
timelike future-complete
timelike past-complete

if any future-directed
past-directed

timelike geodesic $t \mapsto \exp(tX)$

may be defined on $[0, \infty)$

Similarly = lightlike future-complete
past-complete

Note Lemma 2.51 (Bär)

Let M be a semi-Riem. mfd.

$c: [0, b] \rightarrow M$ a geodesic

Then TFAE

(i) $\lim_{t \rightarrow b} c(t)$ exists in M

(ii) $\exists \varepsilon > 0$ s th. c can be extended to a geodesic

$c: [0, b + \varepsilon) \rightarrow M$.

Pf: (ii) \Rightarrow (i), (i) \Rightarrow (ii) Bär

\mathbb{R} Lorentzian
Thus M is not timelike future-
complete

$\Leftrightarrow \exists$ f.d. timelike geodesic

$\gamma: (a, b) \rightarrow M$ such that

$\lim_{t \rightarrow b} \gamma(t)$ does not exist.

Example

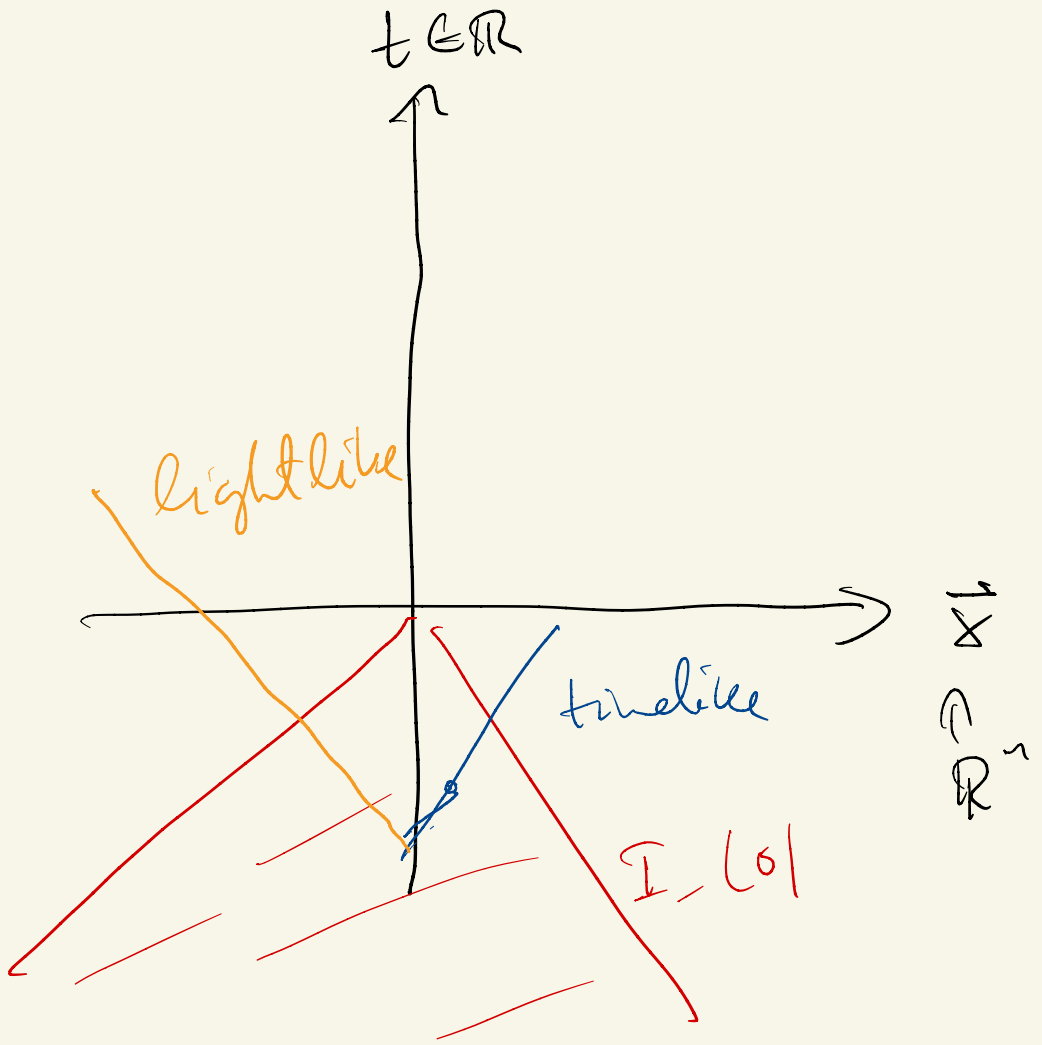
$$M = I_{-}(0) \subset \mathbb{R}^{n,1}$$

not timelike future-complete

not lightlike future-complete

timelike and lightlike

past-complete.



$A \subset M$ $E_{\pm}(A) = J_{\pm}(A) \setminus I_{\pm}(A)$
 event set of A

Def 2.118 A subset $A \subset M$

is called future-trapped if
past

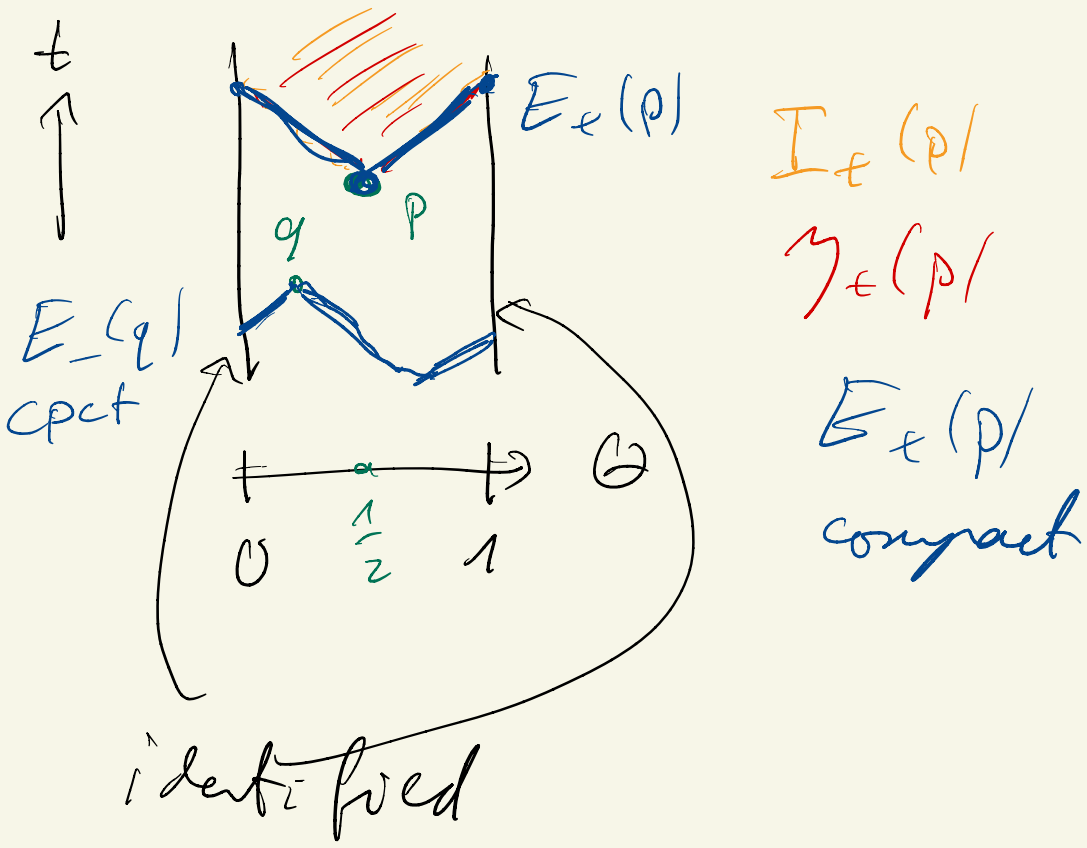
• A is achronal

• A is closed

• if $\boxed{E_{\epsilon}(A) = J_{\epsilon}(A) \setminus I_{\epsilon}(A)}$

$E_{\epsilon}(A)$ is compact.

Example $(M, g) = (\mathbb{R} \times S^1, -dt^2 + d\theta^2)$
 $= \mathbb{R}^{1,1} / \mathbb{Z} e_1$



$A = \{p\}$ is future trapped and past-trapped

Remarks

1.) A acronal $\Leftrightarrow A \cap \bar{I}_\epsilon(A) = \emptyset$

$$\Rightarrow A \subset \underbrace{\underbrace{I_\epsilon(A) \cup \bar{I}_\epsilon(A)}_{\text{cpct}}}_{\text{closed}} = E_\epsilon(A)$$

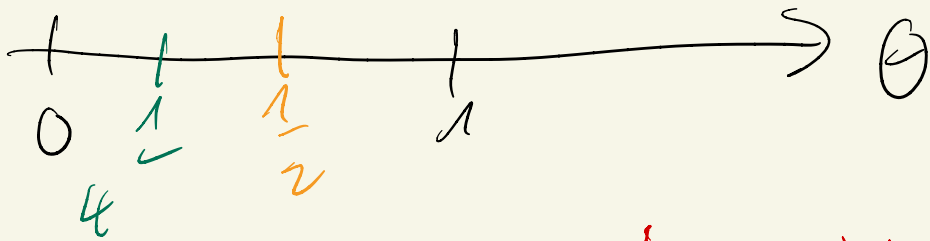
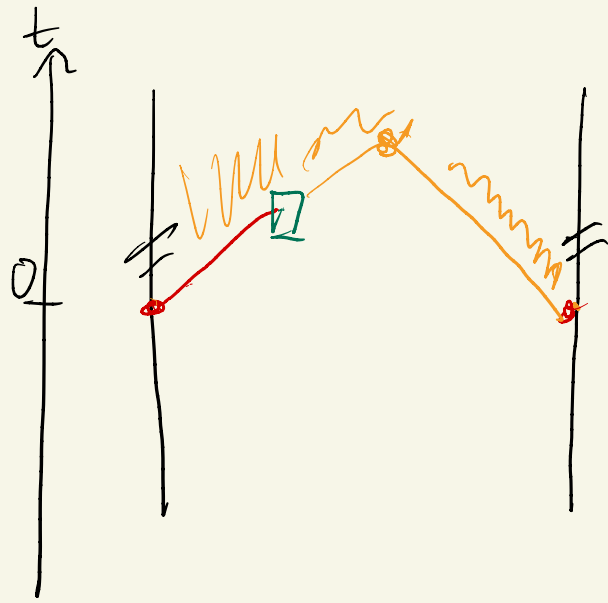
$$\Rightarrow A \text{ cpct}$$

Thy A trapped $\Rightarrow A$ compact.

1a) We cannot remove "closedness" in Def 2.118 without changing the defn.

$$M = \mathbb{R}^{2,1} / \mathbb{Z} e_1$$

$$\bar{A} \equiv \left\{ \begin{pmatrix} t \\ t \bmod 1 \end{pmatrix} \mid 0 \leq t \leq \frac{1}{4} \right\} \subset M$$



$$\gamma_\epsilon(A) = \gamma_\epsilon(\bar{A}) = \left\{ \begin{pmatrix} t \\ x \bmod 2 \end{pmatrix} \mid \begin{array}{l} t \geq \\ |x|, \\ -\frac{1}{2} \leq x \leq \frac{1}{2} \end{array} \right\}$$

$$I_{\epsilon}(A) = I_{\epsilon}(\tilde{A}) = \left\{ \begin{pmatrix} t \\ x \bmod 2\ell \end{pmatrix} \mid \begin{array}{l} t > \\ |x|, \\ -\frac{1}{2} \leq x \leq \frac{1}{2} \end{array} \right\}$$

$$E_{\epsilon}(A) = E_{\epsilon}(\tilde{A}) = \left\{ \begin{pmatrix} t \\ x \bmod 2\ell \end{pmatrix} \mid \begin{array}{l} t = \\ |x|, \\ -\frac{1}{2} \leq x \leq \frac{1}{2} \end{array} \right\}$$

cpdt

2.) For arbitrary $A \subset M$, the subset $E_\varepsilon(A)$ is achronal:

$$p \in \underline{I}_\varepsilon(\underbrace{E_\varepsilon(A)}_{\gamma_\varepsilon(A)}) \subset \underline{I}_\varepsilon(\gamma_\varepsilon(A))$$

$$= \underline{I}_\varepsilon(A) \Rightarrow p \notin E_\varepsilon(A) = \gamma_\varepsilon(A) \setminus \underline{I}_\varepsilon(A)$$

$$\underline{I}_\varepsilon(E_\varepsilon(A)) \cap E_\varepsilon(A) = \emptyset$$

$$\Rightarrow E_\varepsilon(A) \text{ achronal.}$$

Lemma 2.12.1 $m = n + 1$

Let M be an m -dim Lorentzian

mfld, $p \in M$, $l \in T_p M$

lightlike, and e_1, \dots, e_{m-2}

spacelike, orthonormal, $e_j \perp l$

Then $\text{ric}(l, l) = \sum_{j=1}^{m-2} \langle R(l, e_j)e_j, l \rangle$

Proof: Assume $l = \lambda(e_0 + e_{m-1})$

$\lambda \in \mathbb{R}_{>0}$, $\{e_0, e_1, \dots, e_{m-1}\}$ gen-orth.

basis
for $T_p M$

$$\langle e_0, e_0 \rangle = -1$$

e_1, \dots, e_{m-2} basis for $l^\perp \subset T_p M$

$$0 = \langle R(l, l) e_0, l \rangle \quad (1)$$

$$= \lambda (\langle R(l, e_0 + e_{m-1}) e_0, l \rangle)$$

$$0 = \lambda (\langle R(l, e_0 + e_{m-1}) e_{m-1}, l \rangle) \quad (2)$$

$$0 = \frac{(1) - (2)}{\lambda} = \langle R(l, e_0) e_0, l \rangle$$

$$- \langle R(l, e_{m-1}) e_{m-1}, l \rangle$$

\Rightarrow Lemma

$$\langle R(l, x) y, l \rangle = \langle R(l, y) x, l \rangle$$

Prop 2.122 Let M be a Lorentzian
 mfd, $P \subset M$ a spacelike submfd
 of codim 2 and mean curvature
 vector field \vec{H} .

Let $c: [0, b] \rightarrow M$ be lightlike
 geodesic, $c(0) \in P$, $\dot{c}(0) \perp T_{c(0)} P$

Assume (i) $\text{ric}(\dot{c}(t), \dot{c}(t)) \geq 0 \quad \forall t \in [0, b]$

(ii) $\langle \vec{H}(p), \dot{c}(0) \rangle \geq \frac{1}{b}$

Then c has a focal point in $(0, b]$.

$$p \in P, (T_p P)^\perp = N_p P \cong \mathbb{R}^{1,1}$$

Proof: Assume that c has no focal point in $(0, b]$.

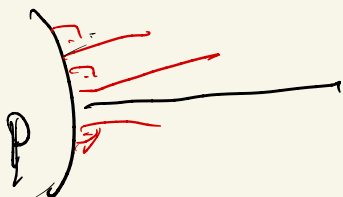
a) Let e_1, \dots, e_{m-2} be an orb of $T_p P$. $\gamma_i \in \mathcal{L}(c^* TM)$ Jacobi field with

$$i \in \{1, \dots, m-2\} \quad \gamma_i(0) = e_i,$$

$$\frac{\nabla \gamma_i}{dt}(0) = -S_c(0)(e_i)$$

$$i=0 \quad \gamma_0(0) = 0, \quad \frac{\nabla \gamma_0}{dt}(0) = \dot{c}(t)$$

$$\gamma_0(t) = t \cdot \dot{c}(t)$$



b) Claim $y_0(t), \dots, y_{m-1}(t)$

is a basis of $(\dot{c}(t))^\perp$

for all $t \in (0, b]$.

$$\langle y_0(t), \dot{c}(t) \rangle = \underbrace{f \langle \dot{c}(t), \dot{c}(t) \rangle}_{=0} = 0$$

$$\stackrel{i>0}{=} \frac{d^2}{dt^2} \langle y_{i-1}, \dot{c} \rangle = \left\langle \frac{D^2}{dt^2} y_{i-1}, \dot{c} \right\rangle$$

$$= \langle R(\dot{c}, y_{i-1}) \dot{c}, \dot{c} \rangle = 0$$

$$\frac{d}{dt} \langle y_i(t), \dot{c}(t) \rangle \Big|_{t=0} = \left\langle \frac{D}{dt} y_i, \dot{c} \right\rangle \Big|_{t=0}$$

$$= - \left\langle \underbrace{S_{\dot{c}(0)}(e_i)}_{\in T_p P}, \underbrace{\dot{c}(0)}_{\perp T_p P} \right\rangle = 0$$

$$\langle \gamma_i(t), \dot{c}(t) \rangle = \langle e_i, \dot{c}(t) \rangle = 0$$

$$\Rightarrow \gamma_i(t) \perp \dot{c}(t) \quad \forall t$$

Lin. independence of $\gamma_0(t), \dots, \gamma_{m-2}(t)$

Assume $t_0 \in (0, b]$

$$\sum_{i=0}^{m-2} \alpha_i \gamma_i(t_0) = 0, \quad \alpha_i \in \mathbb{R}$$

not all are 0.

$$\gamma := \sum_{i=0}^{m-2} \alpha_i \gamma_i, \quad \gamma(t_0) = 0$$

$$\dot{\gamma}(t_0) = \sum_{i=1}^{m-2} \alpha_i e_i \in T_{t_0} P$$

$$\pi^{\text{tan}} \left(\frac{Dy}{dt} (0) \right)$$

$$= \pi^{\text{tan}} \left(\sum_{i=0}^{m-2} \alpha_i \frac{Dy_i}{dt} (0) \right)$$

$$= \pi^{\text{tan}} \left(\sum_{i=0}^{m-2} \alpha_i (-1) S_{\dot{c}(0)}(e_i) \right)$$

$$= -\pi^{\text{tan}} \left(S_{\dot{c}(0)} \left(\underbrace{\sum_{i=0}^{m-2} \alpha_i e_i}_{= y(0)} \right) \right)$$

$$= -\pi^{\text{tan}} \left(S_{\dot{c}(0)}(y(0)) \right)$$

$y \neq 0$ to is a focal point of γ

There are $\gamma_0(t), \dots, \gamma_{m-2}(t)$ lin. ad.